# Progress in Computing Pi, 250 BCE to the Present From Polygons to Elliptic Integrals

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The history of computing  $\pi$  offers a high-level view of several aspects of mathematics and its history. It also says something about aspects of computing and the history of computers. Beyond that, since the practical value of more accurate approximations to  $\pi$  was exhausted by the turn of the 18<sup>th</sup> century and the mathematical value of extremely high-precision approximations is very limited, it says something about human culture.

In this article, I give background for and outline the four fundamentally different techniques (measurement and three methods of calculation) of finding approximations that have been used, and I discuss both the mathematics and the computer science behind computing  $\pi$  to the trillions of decimal places that are now known. I assume mathematical background about the level of what used to be called "college algebra" but is now known as precalculus. Two non-obvious conventions are: (1) Numbers of decimal places or approximate values for  $\pi$  in **boldface italics** indicate values that were world records when they were found. (2) References giving only a page number are to Beckmann [2].

The record for accuracy in computing approximations to  $\pi$  now stands at 12.1 trillion decimal places [16]. There's a great deal of history behind this remarkable feat. Indeed, Beckmann comments in his fascinating book A History of Pi [2], "The history of  $\pi$  is a quaint little mirror of the history of man." (p. 3)<sup>1</sup> But I'll concentrate in this short paper exclusively on its mathematical history, a history that divides into periods—perhaps "phase" is a better term, since the last two overlap—surprisingly well.

Phase 0 (before ca. 250 BCE): Measurement. In earliest times, of course, no method for actually computing  $\pi$  was known. The first values were found by the Egyptians, Babylonians, and others by measuring the circumferences and diameters of circles, and dividing the former by the latter.

<sup>1</sup> My references are to the third (1974) edition of this book. However, based on what the preface to the third edition says, that "edition" is essentially just the second edition with some minor errors corrected. In fact, the title page of an old copy of what is actually the third edition—as indicated by the presence of the third-edition preface—claims it's the second edition! A more recent copy of the alleged third edition simply removes that claim; the title page still does not say it's the third edition, but it includes the same

"third edition" preface as the old copy.

Phase 1 (ca. 250 BCE to 1654): The "method of exhaustion", considering polygons circumscribed and inscribed in a circle and iterating over the number of sides (Archimedes and others). Borwein et al. [4] comment that "despite the long pedigree of the problem, all nonempirical calculations have employed, up to minor variations, only three techniques"; this was the first.

Phase 2 (ca. 1665 to 1990): The discovery of increasingly efficient infinite series and iterative algorithms for computing  $\pi$ , starting with algorithms based on calculus (Newton, John Machin, Euler, and others) and ending with those based on the transformation theory of elliptic integrals (following up work by Gauss, Legendre, and Ramanujan). These are Borwein et al.'s second and third techniques, respectively.

Phase 3 (ca. 1930 to the present, overlapping phase 2): The application of increasingly powerful digital computers (including electromechanical calculators, the immediate predecessors of digital computers) to the problem, using Borwein et al.'s second and third techniques.

Needless to say, not much precision could be obtained by measurement; the most accurate value of Period 0 seems to have been 25/8 = 3.125, apparently known to the Babylonians by around 1800 BCE [14] (though they and other ancient peoples generally used less accurate values), and about 0.01659 too low <sup>2</sup>

Even in the purely mathematical history of computing  $\pi$ , there is far more than can be covered in one article. I'll say no more about the "measurement" phase, but will summarize the other phases, focusing on one or two major events or features of each.

### Infinite Series and Sequences and Rapidity of Convergence

As I have said, the three known families of techniques for computing  $\pi$  all involve infinite series or iterative algorithms. For a sequence (produced by iteration) or a series to be usable for computing anything, it must, of course, converge. For practical use, it must converge fairly rapidly. And to compute something to billions of significant figures, much less the trillions to which  $\pi$  is now known, it must converge extremely rapidly. The history of computing  $\pi$  furnishes extraordinary demonstrations of this fact. To fully appreciate how much more quickly modern  $\pi$ -computing series converge than the earliest ones requires an idea from numerical analysis: order of convergence. Convergence with order 1—linear convergence—means that for each iteration you get the same number of new digits. For example, it might be 10 new correct digits per iteration: after one iteration, 10 correct digits; after two iterations, 20 correct; after three, 30 correct, etc. Convergence with order 2, quadratic convergence, means the number of new correct digits doubles with each iteration. Thus, the first iteration might provide 4 correct digits; the second, 8 new correct digits; the third, 16; and so forth. Cubic convergence means the number of new correct digits triples each time. These are examples of superlinear convergence, where the number of new correct digits per iteration grows without bound. By contrast, there is also *sublinear* convergence, where the number of iterations needed to get one new correct digit grows without bound.

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<sup>&</sup>lt;sup>2</sup> Weisstein comments [14] that one early Egyptian value,  $256/81 \approx 3.160494$ , is more accurate than the Babylonians' 25/8, but that's not true; it's actually slightly less accurate.

To sum up, with *linear* convergence, the amount of computation per digit is constant; with *superlinear* convergence, the computation per digit goes down and down; and with *sublinear* convergence, computation per digit goes up and up. (For technical details, see the Appendix: Rapidity of Convergence.) Here is a striking example of the contrast. The first  $\pi$ -computing series ever found, the Gregory (or Madhava or Leibniz) series, converges so slowly that, if an efficient implementation on the fastest supercomputer now in existence had started running at the Big Bang, it would have computed no more than about 65 decimal places so far<sup>3</sup>. On the other hand, a reasonable implementation of a state-of-the-art series running on an ordinary desktop computer can produce millions of digits of  $\pi$  in under 30 seconds.<sup>4</sup> Each term of the Gregory series, though it can be computed very quickly, gives you very little information—in fact, less and less as you go on; by contrast, each successive term of a modern series, though it takes considerably more computation, gives you more information. In the terminology just discussed, the Gregory series exhibits sublinear convergence; the modern series, superlinear convergence.

## Phase 1. The Method of Exhaustion

Around 250 BCE, Archimedes devised the first way to obtain  $\pi$  to any desired accuracy, given enough patience, by applying the method of exhaustion invented by Eudoxus a century earlier. Archimedes realized that the perimeters of regular polygons circumscribed about and inscribed within a circle give upper and lower bounds for  $\pi$ . The more sides the polygon has, the more closely it approximates the circle—"exhausting" the space between the circle and the polygon, hence the name—and the tighter the bounds. (Strogatz [12] gives an unusually clear and readable description, addressed to the layperson, of both this process and the way Archimedes used it to find the area of the circle.) The results can easily be computed without trigonometry; in the words of [4]:

[Recursively] calculate the length of circumscribed and inscribed regular 6 x  $2^n$ -gons about a circle of diameter 1. Call these quantities  $a_n$  and  $b_n$ , respectively. Then  $a_0 := 2\sqrt{3}$  and  $b_0 := 3$  and, as Gauss' teacher Pfaff discovered in 1800,

$$a_{n+1} = \frac{2a_n b_n}{a_n + b_n}$$
 and  $b_{n+1} = \sqrt{a_{n+1} b_n}$ 

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<sup>&</sup>lt;sup>3</sup> Beckmann [2] (p. 140) cites an estimate that "to obtain 100 decimal places, the number of necessary terms would be no less than  $10^{50}$ !" At first glance, that number seems far too low: the partial sums after  $10^{50}$  and  $10^{50} + 1$  terms differ by about  $10^{-50}$ , i.e., after only about 50 decimal places. However, the arithmetic mean of two consecutive partial sums of the Gregory series is accurate to roughly the product of the last two terms, giving about twice as many correct decimal places. So, for 60 decimal places,  $10^{30}$  terms is a reasonable estimate; for 70 places,  $10^{35}$  terms. Now, the age of the universe is around  $4.3 \times 10^{17}$  sec. As of this writing, the fastest supercomputer in existence does about 34 petaFLOPS, i.e.,  $3.4 \times 10^{16}$  floating-point operations per sec. It's not easy to convert this figure to an estimate of speed evaluating the Gregory series, but assuming the equivalent of 10 floating-point operations per 100-digit term —surely far too low—yields  $3.4 \times 10^{15}$  terms per sec. But even at that rate, fewer than  $1.5 \times 10^{33}$  terms could have been produced by now.

<sup>&</sup>lt;sup>4</sup> A Java version of T. Ooura's program pi\_css5, which is nowhere near the fastest program available [6], running on an 2.4 GHz MacBook found over 4 million places in 27 sec. See the table under Phase 3, below.

This process converges to  $\pi$ , with the bounds  $a_n$  and  $b_n$  tightening by a factor of four on each iteration [3]. Thus, the sequence of means converges linearly, with a rate of convergence of 1/4. With polygons of 96 sides (i.e., n = 4 in the above formulation), Archimedes obtained  $223/71 < \pi < 22/7$ ; in decimal form, roughly  $3.140845 < \pi < 3.142857$  (average  $\approx 3.141873$ ).

For the next 1800 years, others improved on his results simply by using much larger numbers of sides. Among them, in the 5<sup>th</sup> century CE, the Chinese Tsu Ch'ung Chi found 7 *decimal places*; in 1615, after years of effort, the Dutchman Ludolph van Ceulen found 35 *places*, using polygons of no less than  $2^{62}$  (roughly  $4.6 \cdot 10^{18}$ ) sides! A few years later (1621) another Dutchman, Willebrord Snellius, observed that using regular polygons to approximate circles in a different way makes it possible to bring the bounds for a given number of sides much closer together (p. 111). Where Archimedes had gotten only 2 decimal places from polygons of 96 sides, Snellius obtained  $3.1415926272 < \pi < 3.1415928320$ , for 6 places.

## Phase 2. Methods from Calculus and Transformations of Elliptic Integrals

The invention of differential calculus by Newton, Leibniz, and others in the mid- $17^{th}$  century led to the discovery of a long list of infinite series for computing  $\pi$ , all (or almost all) exploiting the Taylor series for trigonometric functions [1, 2]. Much later, the transformation theory of elliptic integrals provided a basis for a set of iterative methods to compute  $\pi$ .

The earliest series for  $\pi$  is a special case of the series for arctangent, known to Indian mathematicians by the early 1500's, but apparently discovered as early as 1400, probably by Madhava of Sangamagrama [7]. This work was unknown in the west until quite recently; the arctangent series was rediscovered in 1671 by James Gregory and a few years later by Leibniz (p. 132–133). The series is:

$$\arctan x = 1 - \frac{x^3}{3} + \frac{x^5}{5} - \frac{x^7}{7} + \dots + \frac{(-1)^n x^{2n+1}}{2n+1} + \dots$$

This is simply the Taylor series for arctangent, but Taylor's work—extending Gregory's—was still a few decades in the future. As Liebniz observed (it is not known if Gregory did), since arctan  $1 = \pi/4$ , substituting 1 for x yields:

$$\frac{\pi}{4} = 1 - \frac{1}{3} + \frac{1}{5} - \frac{1}{7} + \dots + \frac{(-1)^n}{2n+1} + \dots$$

Unfortunately, this elegant series—called by various combinations of the names Madhava, Gregory, and Leibniz, but we'll call it the Gregory formula—converges sublinearly. In fact, as I have already said, it converges so slowly, it's virtually useless for calculation. The number of terms needed to compute n decimal places increases exponentially, specifically as  $10^{n/2}$ , so that adding 12 decimal places (for example) requires increasing the number of terms by a factor of  $10^6$ . But there are many ways to get better performance from the Taylor series.

The "digit hunters" (as [2] calls them) of the 17<sup>th</sup> and 18<sup>th</sup> centuries used the Taylor series, which they employed in different ways to accelerate its convergence, or a series for arctangent discovered by

<sup>5</sup> Some evidence exists that Archimedes went further and obtained considerably more accurate results. See [1], p. 175, or [2], p. 66.

Euler (p. 154). About 1699, the astronomer Abraham Sharp computed 72 decimal places (p. 102) merely by using the Taylor series for  $\pi/6$ , i.e.,  $x = \frac{1}{\sqrt{3}}$ . After factoring out  $\frac{1}{\sqrt{3}}$ , this makes the general term  $\frac{(-1)^n}{3^n \times (2n+1)}$ . This series converges linearly, with a rate of 1/3. But another astronomer,

#### Machin's Formula and Its Derivation

one named Machin, did something much more far reaching.

In 1706, John Machin, a professor of astronomy in London, came up with a strategy that both made the series for  $\pi$  converge much more rapidly than Sharp's approach, and simplified the calculations it required. His formula states:

$$\frac{\pi}{4} = 4 \tan^{-1} \frac{1}{5} - \tan^{-1} \frac{1}{239}$$

Substituting the Taylor series for the arctangents gives:

$$\frac{\pi}{4} = 4\left(\frac{1}{5} - \frac{1}{3 \times 5^3} + \frac{1}{5 \times 5^5} - \dots\right) - \left(\frac{1}{239} - \frac{1}{3 \times 239^3} + \frac{1}{5 \times 239^5} - \dots\right)$$

As Beckmann ([2], p. 145) says, "This was a neat little trick, for the second series converges very rapidly, and the first is well suited for decimal calculations." True. In fact, both series converge linearly, the first with a rate of  $\frac{1}{25}$ , the second with a rate of  $\frac{1}{239^2} = \frac{1}{57,121}$ . But where does Machin's formula come from? Deriving it requires only the tangent double-angle and subtraction identities plus a bit of algebra.

Let  $\tan \beta = 1/5$ . Then we have

$$\tan 2\beta = \frac{2\tan \beta}{1 - (\tan \beta)^2} = \frac{\frac{2}{5}}{1 - \frac{1}{25}} = \frac{\frac{2}{5}}{\frac{24}{25}} = \frac{5}{12}$$

and

$$\tan 4\beta = \frac{2\tan 2\beta}{1 - (\tan 2\beta)^2} = \frac{2(\frac{5}{12})}{1 - (\frac{5}{12})^2} = \frac{\frac{10}{12}}{\frac{119}{144}} = \frac{120}{119}$$

Now,  $\tan (\pi/4) = 1$ , so  $\tan^{-1}(4\beta)$  is already very nearly  $\pi/4$ ; and  $\tan(u - v) = \frac{\tan u - \tan v}{1 - \tan u \tan v}$ Therefore

$$\tan(\frac{\pi}{4} - 4\beta) = \frac{\tan\frac{\pi}{4} - \tan 4\beta}{1 + \tan\frac{\pi}{4}\tan 4\beta}$$

$$= \frac{1 - \tan 4\beta}{1 + \tan 4\beta} = \frac{1 - \frac{120}{119}}{1 + \frac{120}{119}} = \frac{-\frac{1}{119}}{\frac{239}{119}} = -\frac{1}{239}$$

Continuing with some elementary algebra:

$$\tan\left(4\beta - \frac{\pi}{4}\right) = \frac{1}{239}$$
, or, equivalently,  $4\beta - \frac{\pi}{4} = \tan^{-1}\frac{1}{239}$ , hence

$$4\beta - \tan^{-1}\frac{1}{239} = \frac{\pi}{4}$$

Finally, since  $\tan \beta = 1/5$ , we have the desired formula:

$$4\tan^{-1}\frac{1}{5}-\tan^{-1}\frac{1}{239}=\frac{\pi}{4}$$

Machin used his new formula to compute  $\pi$  to 100 decimal places (p. 145), and in 1873, William Shanks used it to find a remarkable *527 places* (p. 103), a record that stood till 1946. (Shanks actually calculated 707 places, but only the first 527 were correct.)

#### **Transformation Theory of Elliptic Integrals**

According to [4], "the third technique, based on the transformation theory of elliptic integrals, provides the algorithms for the most recent set of computations." Elliptic integrals are so called because they were originally studied in connection with finding the arc lengths of ellipses. Since a circle is an ellipse, their connection to  $\pi$  is plausible. Some of these algorithms are based on the early 19th-century work of Gauss and Legendre with the arithmetic/geometric mean; the state-of-the art algorithms come from the more recent work of Srinivasa Ramanujan.

The arithmetic/geometric mean (AGM) is an interesting thing. Unlike the more familiar means—the arithmetic, geometric, and harmonic—there is no closed-form expression for the AGM of two numbers; instead it must be computed iteratively from the arithmetic and geometric means. Work by Gauss and Legendre relating the AGM to the complete elliptic integral led to the formula discovered in 1976 independently by Salamin [11] and Brent:

$$\pi = \frac{\left[2AGM(1, 2^{-1/2})\right]^2}{1 - \sum_{n=1}^{\infty} 2^{n+1} (a_n - b_n)^2}$$

where  $a_n$  and  $b_n$  are the two numbers in the *n*th AGM iteration. Implementations of this formula converge quadratically, i.e., with order 2, so that each iteration doubles the number of correct digits. [3] offers "a mathematically intermediate perspective and some bits of the history" of computing  $\pi$  and various elementary functions using the AGM.

In 1914, Ramanujan published no fewer than 19 new infinite series for  $1/\pi$ ; perhaps the best known is this "amazing sum" [4]:

$$\frac{1}{\pi} = \frac{\sqrt{8}}{9801} \sum_{n=0}^{\infty} \frac{(4n)!}{(n!)^4} \frac{[1103 + 26390n]}{396^{4n}}$$

Following up his work many years later led to series for  $\pi$  that converge with stunning rapidity, series that converge with order 4 (quadrupling the number of correct digits with every term) or order 5 (quintupling the number of correct digits with every term). However, the mathematics involved is far from simple, and discussing it is beyond the scope of this article.

### **Phase 3. Digital Computers**

The advent of digital computers made the ease of base-10 computation of Machin's series irrelevant, but its rapid convergence was advantage enough, and several "Machin-like" formulae dominated the computation of  $\pi$  from the time of its discovery through the 1970's [6]. In the late 1940's, Shanks' record was broken repeatedly by Ferguson and Wrench, who used Machin-like formulae with desk calculators—essentially, electromechanical computers with extremely limited programmability—to find several values, from 710 to 1,120 places. Then fully electronic digital computers became available and things changed dramatically. From a 2,037-place value computed in 1949 by one of the first such computers, the ENIAC (the project was suggested by John von Neumann [10]), the record shot up to 1,001,250 places in 1973, thanks to one of the first supercomputers, the CDC 7600 [6].

Since then, programs based on the transformation theory of elliptic integrals have wiped out all competition, producing first billions and then trillions of digits of  $\pi$ . The first programs of this type used the so-called Gauss-Legendre formula; the more recent programs are based on Ramanujan's work. The current record of *12.1 trillion decimal places* was set in 2013 by Shigeru Kondo and Alexander Yee, with the Ramanujan-based Chudnovsky formula for "main computation" [16]. But the phrase "main computation" implies that some other kind of computation is involved; what is it?

### **Other Factors in Computation**

Computing  $\pi$  to an enormous number of decimal places obviously means doing arithmetic with enormously long numbers. Doing that accurately and efficiently brings up questions of numerical analysis and especially of analysis of algorithms, two areas that are close to the boundary between mathematics and theoretical computer science. Aside from the "main computation" using series that converge incredibly rapidly, a major factor in the speed of the latest generation of  $\pi$ -computing programs is simply that they perform multiplication of numbers billions or trillions of digits long via unusual and extraordinarily efficient methods. Conventional multiplication has a *time complexity* of  $O(n^2)$ ; that is, the number of operations necessary to multiply two numbers of n digits each increases asymptotically as  $n^2$ , so that doubling the number of digits quadruples the number of operations necessary. This is probably what most people would intuitively expect, but it's possible to do a great deal better. The well-known Fast Fourier Transform (FFT) is one example of an algorithm that *can* do much better (Borwein 1989), though it's rarely used for this purpose. Kondo and Yee's 12.1 trillion decimal-place calculation relied on Yee's program "y-cruncher", which is essentially a framework for computing irrational numbers to extremely high precision with great efficiency. Indeed, y-cruncher uses the FFT as well as several other algorithms to do arithmetic [16]. In addition to  $\pi$ , it has been

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<sup>&</sup>lt;sup>6</sup> These series appear in [4], which also comments that convergence with order 5 is close to the theoretical optimum.

used to compute *e* to a trillion digits, and several other constants—the square root and natural log of 2, Apéry's constant, etc.—to many billions of digits.

It seems unlikely that most of the programs based on Machin-like formulae do arithmetic in unusual ways. But writing a program to compute anything very accurately involves other practical considerations, for example, avoiding fractions for as long as possible in order to minimize roundoff error. With Machin's formula, one implication is the need to substitute  $\cot^{-1}(5)$  and  $\cot^{-1}(239)$  for the original  $\tan^{-1}(1/5)$  and  $\tan^{-1}(1/239)$ .

Finally, of course, the programming language and (to a lesser extent) compiler used can make a big difference. Yee's y-cruncher is written in C++, which generally runs very quickly as compared to more recent languages like Java, C-sharp, and Python. A major reason that these "more recent languages" tend to run slowly is that they are generally compiled to instructions for a virtual machine, not for the actual computer hardware; then those instructions are interpreted. In fact, y-cruncher was originally in Java. Though Yee doesn't say, it seems likely the rewrite was motivated primarily by the desire for more speed.

### **Comparison of Three Methods and Programs**

Here's a simple comparison of three ways to compute  $\pi$  and a program for each. The programs are:

- CalcPiGregory: my own straightforward implementation of the Gregory series, written in Java, with Java's standard arithmetic.
- CalcPiMachin: my adaptation of a straightforward implementation of Machin's formula, written in Java, and doing arithmetic via BigDecimal, the standard Java arbitrary-precision arithmetic package.
- pi\_fftcs (also known as pi\_css5): an implementation of the Gauss-Legendre formula by T.Ooura, using FFTs for arithmetic, converted to Java by Hazeghi [5]. Despite its overwhelming increase in efficiency over even Machin's formula, recall that algorithms based on Ramanujan's work can be even faster.

The first two programs are available at http://www.informatics.indiana.edu/donbyrd/Teach/math/.pi fftcs is available at http://www.kurims.kyoto-u.ac.jp/~ooura/pi fft.html.

Timings were done on an 2010-vintage MacBook with a 2.4 GHz Intel Core 2 Duo CPU and 2 GB of RAM, averaging three runs for each.

Method	Terms for 100 dec. places	Terms for 100K dec. places	Implementation	Approx. lines of code	Execution time for 100K places
Madhava- Gregory series	≥10 <sup>50</sup>	10 <sup>50,000</sup> (?)	CalcPiGregory	100	
Machin's formula	77 + 23	71,537 + 21,024	CalcPiMachin: in Java, with Java's standard arbitrary-precision arithmetic	120	40.5 sec.
Gauss-Legendre formula, AGM	3 (176 places)	13 (131,072 places)	pi_fftcs/pi_css5: in Java, with arithmetic via FFT	2,300	1.62 sec.

#### **Conclusions**

The obvious question is: Why would anyone wish to compute thousands or millions of digits of  $\pi$ , much less 12.1 trillion? Certainly not for any ordinary scientific or technical purpose. In the words of the American astronomer Simon Newcomb, "Ten decimal places of  $\pi$  are sufficient to give the circumference of the earth to a fraction of an inch, and thirty decimal places would give the circumference of the visible universe to a quantity imperceptible to the most powerful microscope." (This statement is widely attributed to Newcomb and, as far as I know, never to anyone else; but I've been unable to find a reliable source for it.) Since Newcomb's death in 1909, microscope technology has improved tremendously, and the "visible universe" is larger as well, so nowadays we'd need to work harder to elude the scrutiny of "the most powerful microscope": 40 decimal places should suffice. But, as I have said,  $\pi$  was known to far greater accuracy than that—71 places—by the beginning of the 18<sup>th</sup> century.

The closest thing to a real-world application of vast numbers of digits of  $\pi$  is probably testing newly-developed or overclocked processors. In the words of Wikipedia (2011), "Overclocking is the process of operating a computer component at a higher clock rate (more clock cycles per second) than it was designed for or was specified by the manufacturer. This is practiced more by enthusiasts than professional users seeking an increase in the performance of their computers, as overclocking carries risks of less reliable functioning and damage." The program SuperPi has been popular for this application for years. But it should be pointed out that, while testing such hardware requires an

<sup>&</sup>lt;sup>7</sup> It seems reasonable to equate Newcomb's "visible universe" with what is now called the "observable universe". The diameter of the observable universe is estimated at about 100 billion light years (this is possible despite the fact that only about 13 billion years have elapsed since the Big Bang because expansion of the early universe was not strictly limited by the speed of light). On the other end of the scale, transmission electron microscopy has a resolution limit of around 0.05 nanometers. A light year is about  $9.46 \times 10^{15}$  meters, so  $10^{11}$  light years is about  $9.46 \times 10^{26}$  meters; the ratio of that distance to  $5 \times 10^{-11}$  meters is less than  $10^{38}$ . One could argue that the appropriate "yardstick" at the small end of the scale is not the resolution of any existing microscope but the Planck distance; that brings in another 20 or so orders of magnitude, which still limits the number of decimal places of π needed to only 60 or so. [8] quotes several other estimates of the number of useful places of π.

enormous amount of computation, and while calculating  $\pi$  is a handy example of a process that does that, it is by no means the only task that does so.

On the purely mathematical side, people have been wondering for a long time about issues such as whether  $\pi$  is a *normal* number, or at least a normal number in base 10: that is, whether it has asymptotically equal distribution of digits and of all digit sequences of a given lengths in base 10. (Several papers on various of the record-breaking calculations discuss this point, and it appears to have been Neumann's motivation for the calculation described in [10].) Of course no computation to a finite number of digits can decide questions like this, but they *can* provide hints that might lead eventually to proofs one way or the other. To my knowledge, all evidence so far suggests that  $\pi$  *is* normal. Wagon's paper [13], from 1985, considers only the first 10 million digits; but more recent publications, e.g., [1], seem to agree.

A more detailed and discussion of this topic appears under the heading "Why do we care?" in Offner's very interesting article [8].

## **Challenges for Students**

The expansion of  $\pi$  to a large number of decimal places fascinates many people. While teaching high-school algebra, I once offered extra credit to any student who memorized from 20 to 50 decimals place. Quite a few took me up on the offer, and several clearly enjoyed it. But there are several ways in which computing  $\pi$  could be used in either the mathematics or the computer-programming classroom.

- For elementary calculus, of course, it offers many examples of convergent infinite series, some simple and similar to well-known series, some less so.
- For numerical analysis, the rapidity of convergence of series is an important topic.
- For programming, implementing almost any algorithm in "standard" precision in any language is fairly straightforward. Implementing in arbitrary precision is more challenging, and implementing an algorithm in such a way as to be able to compute, say, billions of decimal places in a reasonable amount of time is considerably more so.

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# Appendix: Rapidity of Convergence

The following discussion refers to sequences, while we are equally interested in infinite series; but we can apply the same terminology simply by considering the sequence of partial sums of each series. The terms below (except perhaps for "sublinear") are standard in numerical analysis; see any textbook, or [15].

The speed at which a convergent sequence approaches its limit is described by its *order of* convergence and, in many cases, its rate of convergence. In particular, suppose that the sequence  $\{x_k\}$  converges to the number L. We say that this sequence **converges linearly** to L if there exists a number  $\mu$  with  $0 < \mu < 1$  such that

$$\lim_{k\to\infty} \frac{|x_{k+1} - L|}{|x_k - L|} = \mu.$$

If we think of  $|x_k - L|$  as the error at term k, the ratio in the above expression is the factor by which the error is changing at that point; the smaller the ratio, the faster the error is disappearing. The number  $\mu$  is called the **rate of convergence**. However, if

- $\mu = 0$ , then we say the sequence **converges superlinearly**;
- $\mu = 1$ , then we say the sequence **converges sublinearly**.

Furthermore, we say that a superlinearly-convergent sequence converges with order q to L for q > 1 if

$$\lim_{k\to\infty}\frac{\mid x_{k+1}-L\mid}{\mid x_k-L\mid^q}=\mu\mid \mu>0\ .$$

In addition, convergence with order

- q = 2 is called quadratic convergence,
- q = 3 is called **cubic convergence**,
- q = 4 is called quartic convergence,
- etc.

Of course, **linear convergence** is convergence with order 1. This is somewhat counterintuitive, since every geometric sequence that converges has linear convergence, but it's linear in that the number of new correct digits per term is constant.

Another term that is sometimes used is **exponential convergence**. Quadratic convergence is closely related to and implies exponential convergence [3].